



Compositionally induced valence-band offset at the grain boundary of polycrystalline chalcopyrites creates a hole barrier

7 Ug DYfggcb UbX 5`YI `Ni b[Yf`

7]U]cb. `5dd`JYX`D\mg]Mg`@YHfg`87z&%%- \$(`f&\$ \$) t/Xc].`%\$"%\$*` #%&% &)' +

J]Yk `cb`]bY. `hcd.##Xl "Xc]"cf[#"\$%\$*` #%&% &)' +

J]Yk `HUV`Y`cZ7 cbH/bfg. `hcd.##gV]U]cb"U]d"cf[#tcbYbh#U]d#ci fbU`#Jd`#` +&%3j Yf1dXZ/Wj

Di V]g\YX`VmiH`Y`5-D`Di V]g\]b[

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Journal of Applied Physics, Vol. 98, No. 10, October 15, 2005
E 100 44

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(Received 15 June 2005; accepted 26 September 2005; published online 15 November 2005)

First-principles calculations of model grain boundaries (GBs) in CuInSe_2

CuInSe_2 has, in addition, a conduction-band offset at the GB/GI interface, attracting electrons to the GBs. These features explain how polycrystalline chalcopyrite solar cells could outperform their crystalline counterparts. © 2005
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Whereas the Si , GaAs form of conventional semi-conductors (Si, GaAs) has poor transport and electronic properties relative to their crystalline counterparts, surpris-

conventional single-crystal CIS. This model¹ of Cu_2S -
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atoms implies a pronounced affect on the metal–Se bonds there; the distance between the surface III layer and the nearest Se layer ($\delta = -0.17 \text{ \AA}$ in both CIS and CGS) represent a very strong inward relaxation of the metal terminated layer.

(ii) ΔE_v : We find (Fig. 1) that in $\Delta E_v(\text{CGS}) = 0.5 \text{ eV}$, whereas $\Delta E_v(\text{CIS}) = (0.2-0.3) \text{ eV}$. This is a consequence of the shorter Cu–Se bond length in CGS relative to CIS, and is reflected by the greater energetic separation of the Cu bands in CGS from its VBM than in CIS (i.e., larger σ -repulsion compared with CIS). As a result, the removal of Cu atoms upon the creation of a polar GB surface lowers the VBM of CGS more than it does in CIS. The different positions of the Cu band in CIS and CGS are reflected also in the existence of a larger crystal-field splitting $\Delta_{\text{CF}} = \epsilon(\Gamma_{5v}) - \epsilon(\Gamma_{4v})$. 9292 find